

# СВОЙСТВА ПЛАЗМЫ С БОЗЕ КОНДЕНСАТОМ ЗАРЯЖЕННЫХ ЧАСТИЦ

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**XII Марковские чтения**

**14 мая, 2014 года**

## Quite unusual results

found in our papers:

A.D., A. Lepidi, G. Piccinelli,

JCAP 0902 (2009) 027; Phys. Rev D, 80  
(2009) 125009; JCAP 08 (2010) 031;

A.D., A. Lepidi Phys.Lett. A375(2011) 3188.

## Similar results but by another method:

G. Gabadadze, R.A. Rosen,

Phys. Lett. B 658 (2008) 266;

JCAP 0810 (2008) 030;

JCAP 1004 (2010) 028.

Textbook formula for screening:

$$U(r) = \frac{Q}{4\pi r} \rightarrow \frac{Q \exp(-m_D r)}{4\pi r},$$

because the time-time component of the photon propagator acquires “mass”:

$$k^2 \rightarrow k^2 + \Pi_{00}(k) = k^2 + m_D^2,$$

where e.g. for relativistic fermions

$$m_D^2 = e^2 \left( T^2/3 + \mu^2/\pi^2 \right).$$

In presence of charged particle condensate the screening is not exponential but power law and oscillating as a function of distance.

We did not publish our work for about half a year, but then found out that an oscillating screening is known for plasma with degenerate fermions and is observed in experiment - Friedel oscillations.

Physics is different but qualitative behavior is the same.

Strangely until recently the effects on screening from condensate of a charged Bose field were not well studied, though it is a textbook problem.

Consider electrically neutral plasma with large electric charge density of fermions compensated by charged bosons. Bosons condense when their chemical potential reaches maximum value:

$$\mu_B = m_B.$$

Otherwise it is impossible to make larger asymmetry between bosons and antibosons.

Equilibrium distribution of condensed boson

$$f_B^C = C \delta^{(3)}(\mathbf{q}) + \frac{1}{\exp [(E - m_B)/T] \pm 1}$$

is a solution of the kinetic equation, it annihilates the collision integral for an arbitrary constant  $C$ .

$f_{eq}$  is always determined by two parameters, either  $T$  and  $\mu$ , or  $T$  and  $C$ , iff  $\mu = m_B$ .

Collision integral:

$$I_{coll} \sim |A_{fi}|^2 \Pi f_f \Pi(1 \pm f_i) - (\textit{inverse})$$

If T-invariance holds, i.e.  $|A_{if}| = |A'_{fi}|$ :

$$I_{coll} \sim [\Pi f_i (1 \pm f_f) - (i \leftrightarrow f)] d\tau.$$

$I_{coll} = 0$  for arbitrary  $T$  and  $C$   
iff  $\mu = m$ .

If T-invariance is broken and  
 $|A_{if}| \neq |A'_{fi}|$ , :

$$I_{coll}[f_{eq}] \sim \Pi f_i (1 \pm f_f) \left[ |A_{fi}|^2 - |A_{if}|^2 \right]$$

This term is surely non-vanishing!

Do equilibrium distributions remain the same in T-broken theory?



Breaking of T-invariance is unobservable if only one reaction channel is open. In this case  $T_{if} = T_{fi}^*$  with time reflected momenta.

$f_B^C$  annihilates collision integral after summation over all relevant processes, due to S-matrix unitarity or CPT and conservation of probability.

Instead of the detailed balance condition there operates “the cyclic balance” condition

Screening properties of medium are expressed through  $f$  which is not necessarily equilibrium one. In calculations neither imaginary time method which may be inconvenient in presence of condensate or out of equilibrium, nor Matsubara-Keldysh technique are used. We started from the quantum equations of motion, solved them with Green's function up to  $e^2$  order, and averaged corresponding operators not only over vacuum but also over “non-empty” medium.

Operator Maxwell equations:

$$\partial_\nu F^{\mu\nu}(x) = \mathcal{J}_B^\mu(x) + \mathcal{J}_F^\mu(x),$$

where bosonic current is

$$\mathcal{J}_B^\mu(x) = -ie[(\phi^\dagger(x)\partial^\mu\phi(x)) - (\partial^\mu\phi^\dagger(x))\phi(x)] + 2e^2 A^\mu(x)|\phi(x)|^2,$$

plus fermionic current:

$$\mathcal{J}_F^\mu(x) = e\bar{\psi}\gamma_\mu\psi.$$

Using equation of motion for quantum operator  $\phi$ :

$$(\partial^2 + m^2)\phi(x) = \mathcal{J}_\phi(x)$$

express  $\phi$  through  $A_\mu$ :

$$\phi(x) = \phi_0(x) + \int d^4y G_B(x-y)\mathcal{J}_\phi(y),$$

$\phi_0$  is free field operator. In the lowest order in  $e$  take  $\phi = \phi_0$  in  $\mathcal{J}_B^\mu(x)$ .

The r.h.s. of the Maxwell equations in  $e^2$  order is linear (but non-local) in  $A_\mu$  and bilinear in  $\phi_0$  and  $\psi_0$ .

Expand free fields as usually:

$$\phi_0(x) = \int d\tilde{q} \left[ a(q) e^{-iqx} + b^\dagger(q) e^{iqx} \right].$$

Average over medium:

$$\begin{aligned} \langle a^\dagger(q) a(q') \rangle &= f_B(E_q) \delta^{(3)}(q - q'), \\ \langle a(q) a^\dagger(q') \rangle &= [1 + f_B(E_p)] \delta^{(3)}(q - q'). \end{aligned}$$

Unity is subtracted, since it is vacuum contribution.

The Fourier transform of the Maxwell equations in plasma is:

$$\left[ k^2 g^{\mu\nu} - k^\mu k^\nu + \Pi^{\mu\nu}(k) \right] A_\nu(k) = \mathcal{J}^\mu(k)$$

where the boson contribution is:

$$\Pi_{\mu\nu}^B(k) = e^2 \int \frac{d^3q}{2(2\pi)^3 E} [f_B(E, \mu) + \bar{f}_B(E, \bar{\mu})] \left[ \frac{l_\mu l_\nu}{l^2 - m^2} + \frac{p_\mu p_\nu}{(p^2 - m^2)} - 2g_{\mu\nu} \right]$$

where  $l = k + q$ ,  $p = k - q$ , and  $E = \sqrt{q^2 + m^2}$ .

Solving Fourier transformed the linear Maxwell equation for  $A_t$ :

$$\Pi_{tt}(0, k) = \frac{e^2}{2\pi^2} \int_0^\infty \frac{dq q^2}{E_B} [f_B(E_B, \mu_B) + \bar{f}_B(E_B, \bar{\mu}_B)] \left[ 1 + \frac{E_B^2}{kq} \ln \left| \frac{2q + k}{2q - k} \right| \right],$$

plus similar contribution from fermions which neutralize the plasma.

This is the well known result for  $\Pi_{tt}$  in order  $e^2$ .

The screened Coulomb potential is the Fourier transform of tt-component of the photon Green's function in medium:

$$U(r) = e^2 \int \frac{d^3k}{(2\pi)^3} \frac{e^{i\mathbf{k}\cdot\mathbf{r}}}{k^2 + \Pi_{tt}(k)} = \frac{e^2}{2\pi^2 r} \int_0^\infty dk k \frac{\sin kr}{k^2 + \Pi_{tt}}.$$

Asymptotics of the potential of charged impurities is determined by the singularities of  $\Pi_{tt}$  in complex  $k$ -plane.



Comment.

Singularities of  $f(z)$ :

$$f(z) = \int_a^b dy g(z, y)$$

in complex  $z$ -plane appear at such  $z$  for which singularities of  $g(z, y)$ , i.e.  $y_c(z)$ , in complex  $y$ -plane coincides with the bounds of integration,  $a$  or  $b$ , or  $y_c(z)$  pinches the contour of integration.

Two types of singularities:

1. Poles of  $[k^2 + \Pi_{tt}(k)]^{-1}$ . E.g. Debye pole. Necessary to check that the position of the poles are at small  $k$ , such that the infrared asymptotics of  $\Pi_{tt}$  is valid.
2. Singularities of  $\Pi_{tt}(k)$ , originating from the pinch of the integration contour in  $q$ -plane by poles of  $f$  and by branch points of  $\log$ .

Without condensate one obtains the usual  $k$ -independent Debye screening:

$$\Pi_{tt}(0, k) = m_D^2$$

originating from a pole at imaginary axis of  $k$ .

With condensate the corrections to  $\Pi_{tt}$  at low  $k$  are infrared singular:

$$\frac{\Delta\Pi_{tt}}{e^2} = \frac{m_B^2 T}{2k} + \frac{C}{(2\pi)^3 m_B} \left( 1 + \frac{4m_B^2}{k^2} \right)$$

Both terms in the r.h.s. appear only if  $\mu = m_B$ .

Instead of exponential the screening becomes power law and oscillating, depending upon parameters,  $m_j$ :

$$\Pi_{tt} = m_0^2 + m_1^3/k + m_2^4/k^2.$$

May this have something to do with confinement  
Recent paper: P. Gaete, E. Spalucci, 0902.0000  
– confinement in Higgs phase.

Contribution from poles in the limit of large  $m_2 r$  but when power law terms are subdominant:

$$U(r)_{pole} = \frac{Q}{4\pi r} \exp(-\sqrt{e/2m_2}r) \times \cos(\sqrt{e/2m_2}r).$$

Oscillating screening is known for **degenerate** fermions - Friedel oscillations. Observed in experiment.

The screening electrons are waves with  $k = k_F$  (from B. Shklovsky).

Comment.

Friedel oscillations are commonly believed to be zero  $T$  phenomenon, because in this case the integral over  $q$  is in finite interval and the singularity in  $k$  appears when log branch point coincides with the upper limit of the integration.

However the "pinch" method works at  $T \neq 0$  and the  $T = 0$  limit can be recovered by summing all the singularities. Non-zero  $T$  corrections, absent in textbooks can be obtained in this way.

Contribution from the integral along imaginary axis is nonzero because  $\Pi_{00}$  contains an odd in  $k$  term. If  $m_2 \neq 0$ , the dominant term is

$$U(r) = -\frac{12Qm_1^3}{\pi^2 e^2 r^6 m_2^8}.$$

If  $T \neq 0$ ,  $\mu = m_B$ , but the condensate is not yet formed, the asymptotic decrease of the potential becomes:

$$U(r) = -\frac{Q}{\pi^2 e^2 r^4 m_1^3} = -\frac{2Q}{\pi^2 e^2 r^4 m_B^2 T}.$$



Contribution from logarithmic cuts (analogous to Friedel oscillations for fermions).

If the first “pinch” (between the poles of  $f(q)$  and logarithmic branch point) dominates:

$$U_1(r) = -\frac{32\pi Q}{e^2 m_B r^2 \ln^2(2\sqrt{2}z)} e^{-z} \sin z ,$$

where  $z = 2r\sqrt{2\pi T m_B}$ .

NB:  $U_1(r)$  is inversely proportional to  $e^2$  and formally vanishes at  $T \rightarrow 0$ , but remains finite if  $\sqrt{T m_B} r \neq 0$ .

All pinches are comparable:

$$U(r) \approx -\frac{3Q}{2e^2 T^2 m_B^3 r^6 \ln^3(\sqrt{8m_B T} r)}.$$

$U \sim T^{-2}$  valid if  $r \ll 1/\sqrt{16\pi T m_B}$ ,  
i.e. if  $T = 0.1\text{K}$  and  $m_B = 1\text{GeV}$  the  
distance should be bounded from above  
as  $r \ll 3 \cdot 10^{-8}$  cm.

Condensation of vector bosons.

$W^\pm$  would condense in the early universe if lepton asymmetry was sufficiently high.

It leads to large electric asymmetry of  $W$ , such that  $\mu_W = m_W$ .

Plasma neutrality was maintained by quarks and leptons.

Vector bosons have additional degrees of freedom, their spin states, and their condensation demonstrates richer possibilities: Depending on the sign of the pairwise spin-spin coupling  $W$ 's would condense either in  $S = 0$  (scalar) state or in  $S = 2$  (ferromagnetic) state.

Magnetic spin-spin interaction through one photon exchange (similar to Breit equation):

$$U_{em}^{spin}(r) = \frac{e^2 \rho^2}{4\pi m_W^2} \left[ \frac{(S_1 \cdot S_2)}{r^3} - \frac{3}{3} \frac{(S_1 \cdot r)(S_2 \cdot r)}{r^5} - \frac{8\pi}{3} (S_1 \cdot S_2) \delta^{(3)}(r) \right].$$

Here  $\rho$  is the ratio of magnetic moment of  $W$  to the standard one.

For  $S$ -wave the energy is shifted by the last term only.

Local quartic self-coupling of  $W$ :

$$U_{4W}^{(spin)} = \frac{e^2}{8m_W^2 \sin^2 \theta_W} (S_1 S_2) \delta^{(3)}(r).$$

The net result  $U_{em} + U_{4W}$  is negative, so  $S = 2$  state is energetically favorable and spontaneous magnetization in the early universe is possible.

## Suppression due to screening.

The  $ij$  component of  $W$  propagator probably remains massless:  $\Pi_{ij} \sim 1/q^2$ . In QED it is true in perturbation theory, while in non-Abelian theories the screening may occur in higher orders of perturbation theory due to infrared singularities. The screening would diminish the long-range ferromagnetic spin-spin coupling while the local  $W^4$  coupling is not screened.

If the propagator is modified, and the wave function of  $W$ -bosons is constant in space, the spin-spin energy shift is:

$$\delta E \sim \int \frac{d^3 q \delta(q)}{(2\pi)^3} \frac{q^2 (S_1 S_2) - (q S_1)(q S_2)}{q^2 + \Pi_{ss}(q)}$$

$\delta E = 0$ , if  $\Pi_{ss} \neq 0$  at  $q = 0$ .



However, the integration over space should be done with an upper limit,  $l$ , equal to the average distance between the  $W$  bosons so instead of  $\delta^{(3)}(q)$ , we obtain:

$$\int_0^l d^3r e^{iqr} = \frac{4\pi}{q^3} [\sin(ql) - ql \cos(ql)].$$

and the energy shift is non-zero:

$$\delta E = -\frac{(S_1 S_2) e^2}{l^3 m_W^2} F(l),$$

$$F(l) = \int_0^\infty \frac{dx [x \sin x + l^2 \Pi_{SS} \cos x]}{x^2 + l^2 \Pi_{SS}(x/l)}.$$

If  $l^2\Pi_{ss}$  is nonnegligible the e.m. part of the spin-spin interaction would be suppressed and the ferromagnet turns into an antiferromagnet. This might happen at  $T$  above the EW phase transition when the Higgs condensate is destroyed and  $m_{W,Z}$  appear as a result of temperature and density corrections and are relatively small.

The quantitative statement depends upon the **(unknown)** modification of the space-space part of the photon propagator in presence of the Bose condensate of charged  $W$  – **a problem to solve.**

Problem of large scale magnetic fields:  
 $B \sim \mu G$  at several kpc. In the intergalactic space the fields are probably 2-3 orders of magnitude weaker, but still non-vanishing  
Dynamo operates only in galaxies.  
Maybe ferromagnetism of  $W$  might create seeds for large scale magnetic fields.

Screening of magnetic fields is connected with the space-space components, which, in the homogeneous and isotropic case is

$$\Pi_{ij} = a(k) \left( \delta_{ij} - \frac{k_i k_j}{k^2} \right) + b(k) \frac{k_i k_j}{k^2}.$$

Multiplying  $\Pi_{ij}$  by  $\delta_{ij}$  and by  $k_i k_j$  we obtain  $b(k) = 0$  and

$$a(k) = \frac{e^2}{32\pi^3} \int \frac{d^3q}{E} (f + \bar{f}) \left[ 2 + \frac{2k^2(4q^2 - k^2)}{4(\tilde{k}\tilde{q})^2 - k^4} \right].$$

where  $\tilde{k}\tilde{q} = kq \cos \theta$ .

If only the condensate term is retained:

$$a^{(C)}(k) = \frac{e^2 C}{8\pi^3 m_B} \equiv e^2 m_C^2.$$

Since  $a(0) = \text{const} \neq 0$ , the magnetic field is exponentially screened. In absence of magnetic monopoles magnetic field can be screened only by currents, **hence plasma with BEC of electrically charged Bose field must be superconductive-** well known result. (Two regimes of superconductivity: weakly coupled Cooper pairs, i.e. BCS or strong coupling BEC regime.)

If  $\mu < m_B$ , then  $\Pi_{ij}(k)$  vanishes as  $k^2$  in the limit  $k \rightarrow 0$ , as expected:

$$a(k) \approx \frac{e^2 k^2}{24\pi^2} \int \frac{dq}{E} (f + \bar{f})$$

and magnetic fields are not screened.

If  $\mu = m_B$ , even without condensate, i.e. at  $C = 0$ ,  $a(k)$  vanishes only as a first power of  $k$ , which leads to unusual screening features.

$a(k)$  is singular in the limit  $m_B = 0$ , since the integral diverges as  $1/q^2$  at the lower limit of integration,  $q = 0$ . Moreover, a singularity at  $k = 0$  exists for massive particles if  $\mu = m_B$ . The singularity comes from the integration region where  $q \sim k$  due to singularity of  $f(q)$  at low  $q$ . So we obtain for  $k \rightarrow 0$ :

$$a^{(sing)}(k) = \frac{e^2 T}{16} k.$$

For small  $k$  this term would dominate over the usual  $k^2$  term and change the screening behavior.



In the transverse gauge,  $k_j A_j = 0$ , the Maxwell equation can be solved as

$$A_i(x) = \int d^3y G(x - y) \mathcal{J}_i(y).$$

The asymptotics of  $G(r)$  at large  $r$  is determined by

$$G(r) = \frac{(-i)}{4\pi^2 r} \int_0^\infty dk k \frac{(e^{ikr} - e^{-ikr})}{k^2 + a(k)}.$$

$a(k)$  may contain odd terms in  $k$ , so the integral along the half real  $k$ -axis cannot be extended to the whole real axis.

**It leads to non-canonical screening terms.**

Since  $a(k) = k^2 + e^2 m_C^2 + e^2 T k / 16$ , the integral can be rewritten as:

$$G(r) = \frac{(-i)}{4\pi^2 r} \int_0^\infty dk k \left( e^{ikr} - e^{-ikr} \right) \frac{(k^2 + e^2 m_C^2 - e^2 T k / 16)}{(k^2 + e^2 m_C^2)^2 - e^4 T^2 k^2 / 256}.$$

The integral of the even part is expressed through the residues of the poles in the complex  $k$ -plane at:

$$k^{(pole)} = \pm i \sqrt{e^2 m_C^2 - \frac{e^4 T^2}{1024}} \pm \frac{e^2 T}{32}.$$

If  $m_C > e^2 T / 32$ , the screened potential would be exponentially cut with superimposed oscillations. For  $e^2 T \ll 32 m_C$ , the Green function takes the form:

$$G(r) \sim \exp(-em_C r) \cos(e^2 r T / 32).$$

In this case the spatial damping scale is much shorter than the oscillation scale. However, for  $eT \sim m_C$  the scales are comparable.

The contour of the integration of odd in  $k$ , part can be closed in upper or lower quadrant of the complex  $k$ -plane. So in addition to the poles in these quadrants the contributions from the integrals over the imaginary axis are to be included. **They produce a power law screening.**

If  $C = 0$ , but  $\mu = m_B$ , then at small  $k$ :  $a(k) \approx e^2 k T / 16$ , so the Green's function drops as:  **$G(r) \sim 8 / (\pi^2 e^2 r^2 T)$** . This is realized when  $r > 1/T$ .

In presence of condensate the Green's function acquires an additional constant term  $e^2 m_C^2$ . In this case the contribution of the integral over the imaginary axis of  $k$  gives  $G \sim T / (16 e^2 \pi^2 r^4 m_C^4)$ .

Changing of the asymptotics of screening signals formation of the condensate.

**THE END**

## Calculation of singularity.

It is convenient to separate the integral into two parts  $0 < q < k/2$  and  $k/2 < q < \infty$ . In the first part we introduce the new integration variable  $x = 2q/k$ , so  $0 < x < 1$ . In the limit of small  $k$  the energy can be expanded as  $E_B \approx m_B + k^2 x^2 / 8m_B$ . At small  $q$  the distribution function is infrared singular:

$$\left[ \exp \left( \frac{E_B - m_B}{T} \right) - 1 \right]^{-1} \approx \frac{2m_B T}{q^2} = \frac{8m_B T}{k^2 x^2}.$$

Usually this singularity is not dangerous because it is canceled by the integration measure,  $\sim q^2$ . However, the logarithmic term behaves as  $k/q$  for  $q > k$  and as  $q/k$  for  $q < k$ . Thus the integral is finite, but it does not vanish as  $k^2$  when  $k \rightarrow 0$ .

The first part of the integral with  $q < k/2$  can be taken analytically and we obtain:

$$a_1^{(s)}(k) = \frac{e^2 k T}{8\pi^2} \int_0^1 dx \left[ 2 - \left( x - \frac{1}{x} \right) \ln \left| \frac{1+x}{1-x} \right| \right]$$

$$\frac{e^2 k T}{8\pi^2} \left( 1 + \frac{\pi^2}{4} \right).$$

There is also another contribution coming from the part of the integral with  $q > k/2$ . As  $k \rightarrow 0$ , the second part of the integral,  $k/2 < q < \infty$ , gives:

$$a_2^{(s)}(k) = \frac{e^2 k T}{8\pi^2} \int_1^\infty dx \left[ 2 - \left( x - \frac{1}{x} \right) \ln \left| \frac{1+x}{1-x} \right| \right] \\ \frac{e^2 k T}{8\pi^2} \left( -1 + \frac{\pi^2}{4} \right),$$

such that the total contribution is:

$$a^{(s)}(k) = a_1^{(s)}(k) + a_2^{(s)}(k) = \frac{e^2 T}{16} k. \quad (4)$$

For small  $k$  this term could dominate over the usual  $k^2$  term and would change the screening behavior.



so we present the denominator as half of sum and difference of even and odd function as following:

$$f(k) = [f(k) + f(-k)]/2 + [f(k) - f(-k)]/2 \quad (5)$$

Since  $a(k) = k^2 + e^2 m_C^2 + e^2 T k / 16$ , eq. (??) can be rewritten as:

$$G(r) = \frac{(-i)}{4\pi^2 r} \int_0^\infty dk k \frac{\left( e^{ikr} - e^{-ikr} \right) \left( k^2 + e^2 m_C^2 - e^2 T k / 16 \right)}{\left( k^2 + e^2 m_C^2 \right)^2 - e^4 T^2 k^2 / 256}$$

The integral of the even part may be transformed, as usually, into the integral along the whole real axis and after closing the contour in the upper (for  $e^{ikr}$ ) or lower (for  $e^{-ikr}$ ) half-plane we express the result through the residues in the corresponding poles in the complex  $k$ -plane at:

$$k^{(pole)} = \pm i \sqrt{e^2 m_C^2 - \frac{e^4 T^2}{1024}} \pm \frac{e^2 T}{32}. \quad (7)$$

If  $m_C > e^2 T/32$ , the resulting screened potential would be exponentially cut with superimposed oscillations. For  $e^2 T \ll 32 m_C$ , the Green function takes the form:

$$G(r) \sim \exp(-em_C r) \cos(e^2 r T/32). \quad (8)$$

In this case the spatial damping scale is much shorter than the oscillation scale. However, if  $eT \sim m_C$  the scales are comparab